Ionosphere Scintillation Mapping

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Contents

- Calculation technique
- Input Data
- Algorithm
- Examples of results
- Conclusion



Calculation Technique

Kriging technique

- Well suited to random data processing
- Accuracy increases with the number of data
- Allows to produce a map and the corresponding error map







Input Data







Monitor Network 50 Hz receivers



Scintillation indices at 50 Hz (top panel) vs 1 Hz (IGS data: bottom panel)



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Modelling (GISM) MPS model

The model includes an orbit generator (GPS, Glonass, Galileo, ...) <u>Inputs</u>

Geophysical Parameters Scenario

Intermediate calculation : LOS, Ionisation along the LOS

<u>Outputs</u>

Time series, Scintillation indices, Correlation Distances Scattering function

Béniguel Y., P. Hamel, "A Global Ionosphere Scintillation Propagation Model for Equatorial Regions", Journal of Space Weather Space Climate, 1, (2011), doi: 10.1051/swsc/2011004







Modelling vs Measurements (Phase)

Measurements —

The phase RMS value is slightly lower than the S4 value

Some samples exhibit high values (both measurements and modelling) due to the phase jumps

Modelling



Sigma Phi all satellites

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Comparison with Measurements Year 2011 / Guiana / Novatel Measurements







Kriging Algorithm





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Kriging Random Variables

Y the observable, f a vector of factors (location, ...)

Y is the sum of a mean value and a random part

$$Y(f, \omega) = m(f) + \Gamma(f, \omega)$$

$$m(f) = \sum_{1}^{M} a_{i} \varphi_{i}(f)$$

Covariance function

$$c(f_i, f_j) = \langle \Gamma(f_i, \omega) \Gamma(f_j, \omega) \rangle = c(h)$$





Covariance Function

o Many data points

Computation of the experimental covariance

• Few data points

Choice of the covariance function among known models.



Kriging Interpolation Technique

A linear estimator is built with all data points

$$\mathbf{Y}^* = \sum_{1}^{N} \lambda_j \mathbf{Y}(\mathbf{f}_j, \boldsymbol{\omega})$$

The outcome at f_0 is linear

$$Y^{*}(f_{0}) = \sum_{1}^{N} \lambda_{j} (f_{0}, f_{j}, c) Y(f_{j})$$

The estimator gives an exact value at the data points (Lagrange multipliers)





Variance Minimization

• Minimizing the variance
$$|Y^* - Y_0|^2$$

• Taking the inverse of the covariance matrix

$$\mathbf{Y}^*(\mathbf{f}_0) = \mathbf{w}^{\mathrm{T}} \mathbf{Y}$$

- It is an exact interpolator : no errors at data points
- Gives an estimate of the accuracy of the Kriged estimates

$$\operatorname{Min}\left\langle \left| \mathbf{Y}^{*} - \mathbf{Y} \right|^{2} \right\rangle = \sigma_{\operatorname{Min}}^{2}$$



Variogram Estimation





$$\mathbf{h} = \left| \, \bar{\mathbf{r}}_{\mathbf{i}} \, - \, \bar{\mathbf{r}}_{\mathbf{j}} \right|$$



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Variograms Estimated on a Reduced Time Duration

Experimental variograms based on GISM calculations







Covariance Function

$$C(h) = T - \gamma(h)$$
 with $T = 7.410^{-4}$









Cut at Latitude 6°







Cut at Latitude - 3°

Map of S4:kriged data





Cut at Latitude – 3°

0.46

0.44

0.42

0.4

0.38

0.36

0.34

UPMC

-0.8

-0.6

-0.4

SigmaPhi





25

0.45

0.4

0.35

UPMC

-1

-0.8

-0.6

-0.4

-0.2

0.2

0

Latitude

0.4

0.6

0.8

1

Cut at Longitude - 44°





Cut at Longitude - 44°







Cut at Longitude - 36°







Cut at Longitude - 36°

Coupe de SigmaPhi pour une longitude de -36°



Kriging Over a Whole Year



Lima 2012 Dip Lat. 0°65 The covariance function used was deduced from GISM calculations



Conclusion

Analysis of scintillations using the IGS network

(Comparison to 50 Hz data)

Improve the calculation of the variogram (Time as an additionnal factor)

Produce the error maps





